# Markscheme 

November 2017

Physics

Higher level

Paper 3

This markscheme is the property of the International
Baccalaureate and must not be reproduced or distributed to any other person without the authorization of the IB Global Centre, Cardiff.

## Section A

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1. | a |  | single smooth curve passing through all data points $\checkmark$ | Do not accept straight lines joining the dots Curve must touch some part of every $x$ | 1 |
| 1. | b | i | tangent drawn at $80^{\circ} \mathrm{C} \checkmark$ <br> gradient values separated by minimum of $20^{\circ} \mathrm{C} \checkmark$ $9.0 \times 10^{-4}<\mathrm{kJ} \mathrm{~kg}^{-1} \mathrm{~K}^{-2} » \checkmark$ | Do not accept tangent unless "ruler" straight. <br> Tangent line must be touching the curve drawn for MP1 to be awarded. <br> Accept values between $7.0 \times 10^{-4}$ and $10 \times 10^{-4}$. <br> Accept working in J, giving 0.7 to 1.0 | 3 |
| 1. | b | ii | $\mathrm{kJkg}^{-1} \mathrm{~K}^{-2} \checkmark$ |  | 1 |
| 1. | c | i | «0.1×4.198×10=» 4.198 «kJ» or 4198 «J» $\checkmark$ | Accept values between 4.19 and 4.21 | 1 |
| 1. | c | ii | percentage uncertainty in $\Delta T=10 \%$ «2\% + 5\% + 10\% =» 17\% <br> absolute uncertainty $=$ « $0.17 \times 4.198=$ » 0.7 «kJ» therefore 2 sig figs <br> OR <br> absolute uncertainty to more than 1 sig fig and consistent final answer $\checkmark$ | Allow fractional uncertainties in MP1 and MP2 <br> Watch for ECF from (c)(i) <br> Watch for ECF from MP1 <br> Watch for ECF from MP2 <br> Do not accept an answer without justification | 3 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 2. | a | $\begin{aligned} & « \varepsilon=I R+I r » \\ & \frac{1}{I}=\frac{R}{\varepsilon}+\frac{r}{\varepsilon} \checkmark \end{aligned}$ <br> identifies equation with $y=m x+c \checkmark$ «hence $m=\frac{1}{\varepsilon}$ " | No mark for stating data booklet equation <br> Do not accept working where $r$ is ignored or $\varepsilon=I R$ is used <br> OWTTE | 2 |
| 2. | b | «-» $r$ d | Allow answer in words | 1 |


| 3. | $\mathbf{a}$ | «to reduce» random errors $\checkmark$ <br> to reduce absolute uncertainty $\checkmark$ <br> to improve precision $\checkmark$ | OWTTE <br> Do not accept just "to find an average" or just "reduce error" <br> Ignore any mention to accuracy |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{3 .}$ | $\mathbf{b}$ | as the literature value is within the range «9.7-11.1» <br> hence it is accurate $\checkmark$ | OWTTE <br> 1 max |

## Section B

## Option A - Relativity

| Question |  | Answers | Total |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 4. |  |  | light is an EM wave $\checkmark$ <br> speed of light is independent of the source/observer $\checkmark$ | Notes |


| 5. | a |  | a co-ordinate system in which measurements «of distance and time» can be made | Ignore any mention to inertial reference frame. | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5. | b |  | closing speed $=c \checkmark$ |  | 2 |
|  |  |  | 2 «s» $\downarrow$ |  |  |
| 5. | c |  | $u$ and $v$ are velocities with respect to the same frame of reference/Earth $\boldsymbol{A N D} u^{\prime}$ the relative velocity | Accept 0.4c and 0.6c for $u$ and $v$ | 1 |
| 5. | d |  | $\frac{-0.4-0.6}{1+0.24}$ |  | 2 |
|  |  |  | «-»0.81c $\downarrow$ |  |  |
| 5. | e | i | $\gamma=1.25 \checkmark$ |  | 2 |
|  |  |  | so the time is $t=1.6$ «s» $\checkmark$ |  |  |
| 5. | e | ii | gamma is smaller for $B \checkmark$ |  | 2 |
|  |  |  | so time is greater than for $\mathrm{A} \checkmark$ |  |  |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 6. | a | the length of an object in its rest frame <br> OR <br> the length of an object measured when at rest relative to the observer $\checkmark$ |  | 1 |
| 6. | b | world lines for front and back of tunnel parallel to ct axis $\checkmark$ world lines for front and back of train $\checkmark$ which are parallel to $c t^{\prime}$ axis $\checkmark$ |  | 3 |
| 6. | c | $\begin{aligned} & \text { realizes that gamma }=1.25 \checkmark \\ & 0.6 c \checkmark \end{aligned}$ |  | 2 |

(continued...)
(Question 6 continued)

| Question |  | Answers | ALTERNATIVE 1 <br> indicates the two simultaneous events for $t$ frame $\checkmark$ <br> marks on the diagram the different times «for both spacetime points» on <br> the $c t^{\prime}$ axis «shown as $\Delta t^{\prime}$ on each diagram» $\checkmark$ |
| :--- | :--- | :--- | :--- | :--- |
| ALTERNATIVE 2: (no diagram reference) |  |  |  |
| the two events occur at different points in space $\checkmark$ |  |  |  |
| statement that the two events are not simultaneous in the $t^{\prime}$ frame $\checkmark$ |  |  |  |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 7. | a | $\Lambda$ momentum $=900 \checkmark$ |  |  |
|  |  | $\begin{aligned} & E_{\text {proton }}=« \sqrt{p c^{2}+\left(m c^{2}\right)^{2}}=\sqrt{630^{2}+938^{2}}=» 1130 « \mathrm{MeV} » \checkmark \\ & E_{\text {pion }}=« \sqrt{270^{2}+140^{2}}=» 304 « \mathrm{MeV} » \checkmark \\ & \text { so rest mass of } \Lambda=« \sqrt{(1130+304)^{2}-900^{2}}=» 1116 « \mathrm{MeV} \mathrm{c}^{-2} » \checkmark \end{aligned}$ |  | 4 |
| 7. | b | $« E=\gamma m c^{2} \text { so» } \gamma=« \frac{1434}{1116}=» 1.28$ |  | 2 |
|  |  | to give $0.64{ }^{\text {c }} \checkmark$ |  |  |



Option B — Engineering physics

| Question |  | Answers | Total |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 9. | a |  | weight, normal reaction and friction in correct direction $\checkmark$ |
| correct points of application for at least two correct forces $\checkmark$ | Labelled on diagram. |  |  |

(continued...)
(Question 9 continued)

| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 9. | b | ALTERNATIVE 1 $\begin{aligned} & m a=m g \sin \theta-F_{\mathrm{f}} \\ & I \alpha=F_{\mathrm{f}} \times r \end{aligned}$ <br> OR $m r \alpha=F_{f} \checkmark$ $\alpha=\frac{a}{r} \checkmark$ $m a=m g \sin \theta-m r \frac{a}{r} \rightarrow 2 a=g \sin \theta$ <br> ALTERNATIVE 2 $m g h=\frac{1}{2} I \omega^{2}+\frac{1}{2} m v^{2}$ <br> substituting $\omega=\frac{v}{r}$ «giving $v=\sqrt{g h}$ " <br> correct use of a kinematic equation $\checkmark$ <br> use of trigonometry to relate displacement and height «s=hsin $\theta$ » $\checkmark$ | Can be in any order <br> Accept answers using the parallel axis theorem (with $I=2 m r^{2}$ ) only if clear and explicit mention that the only torque is from the weight <br> For alternative 2, MP3 and MP4 can only be awarded if the previous marking points are present | 4 |

(Question 9 continued)

| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 9. | c | 1.68 « $\mathrm{ms}^{-2}$ 》 $\checkmark$ |  | 1 |
| 9. | d | ALTERNATIVE 1 $\begin{aligned} & N=m g \cos \theta \checkmark \\ & F_{\mathrm{f}} \leq \mu m g \cos \theta \end{aligned}$ <br> ALTERNATIVE 2 $\begin{aligned} & F_{\mathrm{f}}=m a « \text { «rom } 9(\mathrm{~b}) » \checkmark \\ & \text { so } F_{\mathrm{f}}=\frac{m g \sin \theta}{2} \checkmark \end{aligned}$ |  | 2 |
| 9. | e | $\begin{aligned} & F_{\mathrm{f}}=\mu m g \cos \theta \\ & \frac{m g \sin \theta}{2}=m g \sin \theta-\mu m g \cos \theta \end{aligned}$ <br> OR $m g \frac{\sin \theta}{2}=\mu m g \cos \theta$ <br> algebraic manipulation to reach $\tan \theta=2 \mu \checkmark$ |  | 3 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 10. | a | $\begin{aligned} & 500000 \times\left(2 \times 10^{-3}\right)^{\frac{5}{3}}=100000 \times V^{\frac{5}{3}} \\ & V=5.25 \times 10^{-3} « \mathrm{~m}^{3} » \end{aligned}$ | Look carefully for correct use of $\mathrm{p}^{\gamma}=$ constant | 2 |
| 10. | b | correct vertical and horizontal lines $\checkmark$ curve between $B$ and $C \checkmark$ | Allow tolerance $\pm 1$ square for $A, B$ and $C$ <br> Allow ECF for MP2 <br> Points do not need to be labelled for marking points to be awarded | 2 |
| 10. | C | use of $P V=n R T \quad O R$ use of $\frac{P}{T}=$ constant $\checkmark$ $T=« 5 \times 290=» 1450$ «K» $\downarrow$ |  | 2 |

(continued...)
(Question 10 continued)

| Question |  | Answers | Notes <br> $\mathbf{1 0 .}$ <br> d <br> area enclosed $\checkmark$ <br> work is done by the gas during expansion <br> OR <br> work is done on the gas during compression $\checkmark$ <br> the area under the expansion is greater than the area under the compression $\checkmark$ |  |  |
| :--- | :--- | :--- | :--- | :--- | :---: |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 11. | a | i | $\begin{aligned} & \text { density }=785 \text { « } \mathrm{kgm}^{-3} » \checkmark \\ & \text { « } \frac{4}{3} \pi(0.03)^{3} \times 785 \times 9.8=» 0.87 \text { «N» } \end{aligned}$ | Accept answer in the range 784 to 786 | 2 |
| 11. | a | ii | $\begin{aligned} & \frac{0.87}{\frac{4}{3} \pi(0.03)^{3} \times 1080 \times 9.8} \\ & \text { OR } \\ & \frac{0.87}{1080 \times 1.13 \times 10^{-4}} \\ & \text { OR } \\ & \frac{785}{1080} \checkmark \\ & 0.727 \text { or } 73 \% \checkmark \end{aligned}$ | Allow ECF from (a)(i) | 2 |
| 11. | b |  | use of drag force to obtain $\frac{4}{3} \pi r^{3} \times 0.04 \times g=6 \times \pi \times 0.0011 \times r \times v \quad \checkmark$ $v=0.071<\mathrm{ms}^{-1} » \checkmark$ |  | 2 |



Option C - Imaging

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 13. | a | i | with object placed between lens and focus <br> two rays correctly drawn $\checkmark$ | Backwards extrapolation of refracted rays can be dashes or solid lines <br> Do not penalize extrapolated rays which would meet beyond the edge of page <br> Image need not be shown | 2 |
| 13. | a | ii | «just less than» the focal length or $f \checkmark$ |  | 1 |
| 13. | b | i | $\frac{1}{10}+\frac{1}{v}=\frac{1}{2}$ |  | 2 |
|  |  |  | $v=2.5$ «m» $\downarrow$ |  |  |
| 13. | b | ii | real, smaller, inverted $\checkmark$ | All three required - OWTTE | 1 |

(continued...)
(Question 13 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 13. | c | i | two correct rays coming from Q $\checkmark$ | Allow any two of the three conventional rays. |  |
|  |  |  | locating $\mathbf{Q}^{\prime}$ below the main axis $\boldsymbol{A N D}$ beyond $f$ to the right of lens AND at intercept of rays $\checkmark$ |  | 2 |
| 13. | c | ii | $\frac{h}{h^{\prime}}=\frac{-x}{x^{\prime}}$ <br> OR $2.5 \text { or } 10 \times 0.3 \text { « m» } \checkmark$ |  | 2 |
|  |  |  | «-» 0.075 «m» $\checkmark$ |  |  |
| 13. | c | iii | towards Q $\downarrow$ | Accept move to the left | 1 |
| 13. | c | iv | spherical aberration $\checkmark$ |  |  |
|  |  |  | top of the shape «R" is far from axis so no paraxial rays $\checkmark$ | For MP2 accept rays far from the centre converge at different points | 2 |



| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 15. | a | realization that $\theta \mathrm{min}$ is the critical angle $\checkmark$ |  | 3 |
|  |  | $\theta=« \sin ^{-1} \frac{1.48}{1.5}=» 80.6 «^{\circ} » \checkmark$ | Accept 1.4 rad Accept 0.16 rad |  |
|  |  | $\beta=« 90-80.6=» 9.4 «^{\circ} » \checkmark$ |  |  |
| 15. | b | because the critical angle is nearly $90^{\circ} \checkmark$ |  |  |
|  |  | then only rays that are «almost» parallel to the fibre pass down it $\checkmark$ |  | 3 |
|  |  | so pulse broadening is reduced $\checkmark$ | OWTTE |  |


| 16. | a | evidence of finding the gradient $\checkmark$ | 2 |
| :---: | :---: | :---: | :---: |
|  |  | $\mu=$ « - gradient $=» 59.9$ « $\mathrm{cm}^{-1} » \checkmark$ |  |
| 16. | b | $I=\frac{I_{0}}{25000} \checkmark$ | 2 |
|  |  | «ln25000 $=\mu x » x=0.17$ «cm» or 1.7 «mm» $\checkmark$ |  |

## Option D - Astrophysics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 17. | a | i | «nuclear» fusion $\checkmark$ | Do not accept "burning" | 1 |
| 17. | a | ii | brightness depends on luminosity and distance/ $b=\frac{L}{4 \pi d^{2}}$ <br> Vega is much further away but has a larger luminosity $\checkmark$ | Accept answer in terms of Jupiter for MP2 | 2 |
| 17. | b | i | a group of stars forming a pattern on the sky AND not necessarily close in distance to each other | OWTTE | 1 |
| 17. | b | ii | the star's position is observed at two times, six months apart, relative to distant stars parallax angle is half the angle of shift $\checkmark$ <br> 2 postions of Earth <br> 6 months apart | Answers may be given in diagram form, so allow the marking points if clearly drawn | 2 |

(continued...)
(Question 17 continued)

| Question |  | Answers | Notes | Total |  |
| :--- | :--- | :--- | :--- | :--- | :---: |
| 17. | b | iii | 1 <br> $\frac{1}{0.13}=7.7 « \mathrm{pc} » \checkmark$ <br> so $d=7.7 \times 3.26=25.1$ «ly» $\checkmark$ |  |  |


| 18. | a | two stars orbiting a common centre «of mass» $\checkmark$ | Do not accept "stars which orbit each other" | $\mathbf{1}$ |  |
| :--- | :--- | :--- | :--- | :--- | :---: |
| 18. | b | « $\lambda \times T=2.9 \times 10^{-3} »$ <br> $T=\frac{2.9 \times 10^{-3}}{115 \times 10^{-9}}=25217$ «K» $\checkmark$ |  |  |  |
| 18. | c |  | use of the mass-luminosity relationship or $\left(\frac{M_{\text {Sirius }}}{M_{\text {Sun }}}\right)^{3.5}=1 \checkmark$ <br> if Sirius $B$ is on the main sequence then $\left(\frac{L_{\text {siriusB }}}{L_{\text {sun }}}\right)=1$ «which it is not» $\checkmark$ | Allow reverse argument beginning with <br> luminosity | 2 |

(continued...)
(Question 18 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 18. | d | i | $\begin{aligned} & \left(\frac{L_{\text {sirusB }}}{L_{\text {Sun }}}\right)=0.025 \checkmark \\ & r_{\text {sirius }}=\text { « } \sqrt{0.025 \times\left(\frac{5800}{25000}\right)^{4}}=» 0.0085 r_{\text {sun }} \checkmark \end{aligned}$ |  | 2 |
| 18. | d | ii | white dwarf $\checkmark$ |  | 1 |
| 18. | e | i | Sirius $A$ on the main sequence above and to the left of the Sun AND Sirius B on white dwarf area as shown $\checkmark$ | Both positions must be labelled <br> Allow the position anywhere within the limits shown. | 1 |

(continued...)
(Question 18 continued)


| 19. | $\mathbf{a}$ | galaxies are moving away <br> OR <br> space «between galaxies» is expanding $\checkmark$ | Do not accept just red-shift |
| :--- | :--- | :--- | :--- | :--- |
| 19. | $\mathbf{b}$ | « $\frac{\Delta \lambda}{\lambda}=» \frac{1.04}{115}=\frac{v}{c} \checkmark$ <br> $0.009 \mathrm{c} \checkmark$ | Accept $2.7 \times 10^{6}$ «m $\mathrm{s}^{-1} »$ <br> Award $[0]$ if 116 is used for $\lambda$ |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 20. | a | interstellar gas/dust «from earlier supernova» $\checkmark$ <br> gravitational attraction between particles <br> if the mass is greater than the Jean's mass $/ M_{j}$ the interstellar gas coalesces $\checkmark$ <br> as gas collapses temperature increases leading to nuclear fusion $\checkmark$ | MP3 can be expressed in terms of potential and kinetic energy | 4 |
| 20. | b | fluctuations in CMB due to differences in temperature/mass/density during the inflationary period/epoch/early universe leading to the formation of galaxies/stars/structures $\checkmark$ gravitational interaction between galaxies can lead to collision $\checkmark$ |  | 3 max |

(continued...)
(Question 20 continued)

| Question |  | Answers | Notes | Total |
| :--- | :--- | :--- | :--- | :--- |
| 20. |  | ALTERNATIVE 1 <br> kinetic energy of galaxy $\frac{1}{2} m v^{2}=\frac{1}{2} m H^{2} r^{2}$ «uses Hubble's law» $\checkmark$ <br> potential energy $=\frac{G M m}{r}=G \frac{4}{3} \pi r^{3} \rho \frac{m}{r}$ «introduces density» $\checkmark$ <br> KE $=$ PE to get expression for critical $\rho \checkmark$ <br> ALTERNATIVE 2 <br> escape velocity of distant galaxy $v=\sqrt{\frac{2 G M}{r}} \checkmark$ <br> where $H_{0} r=\sqrt{\frac{2 G M}{r}} \checkmark$ <br> substitutes $M=\frac{4}{3} \pi r^{3} \rho$ to get result $\checkmark$ |  |  |

